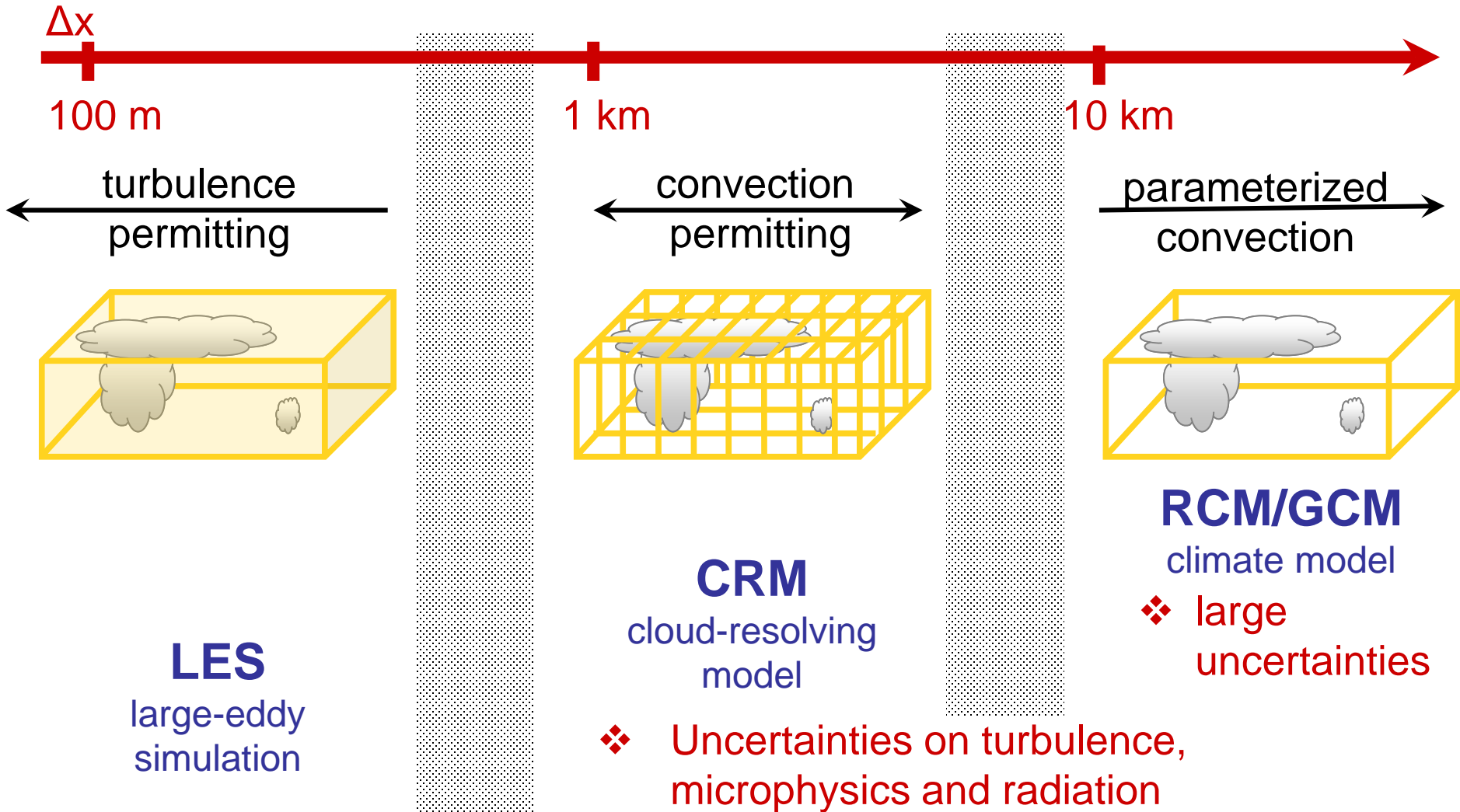


Successes and challenges in the simulation of tropical deep convection at high resolution

Jean-Pierre CHABOUREAU, Thibaut DAUHUT,
Keun-Ok LEE and Irene REINARES MARTINEZ

Laboratoire d'Aérodynamique, University of Toulouse, CNRS, UPS, France

Representation of deep convection



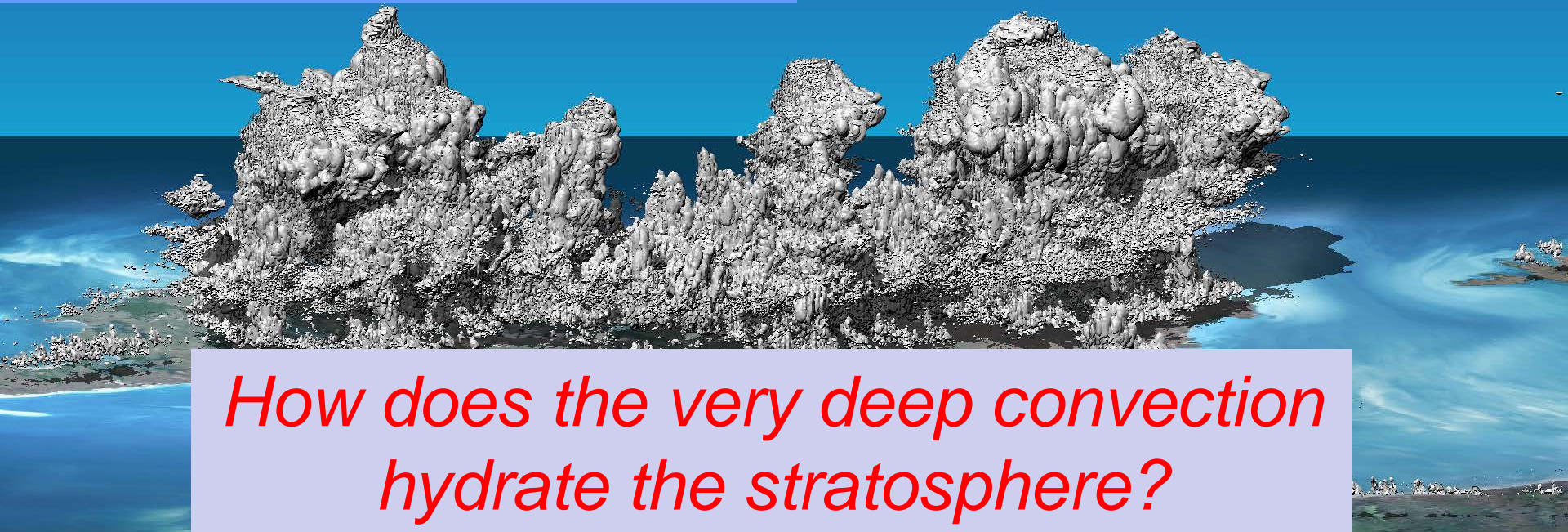
- ❖ Uncertainties on microphysics and radiation

Very deep convection in a Giga-LES



Hector the Convecteur

- 2560 x 2048 x 256, 1.34 billion gridpoints
 $\Delta x=100$ m and $\Delta z=40 - 100$ m
- 10-h simulation on IBM BlueGene-Q
8 million CPU h, 16 kcores, 20 Tb data

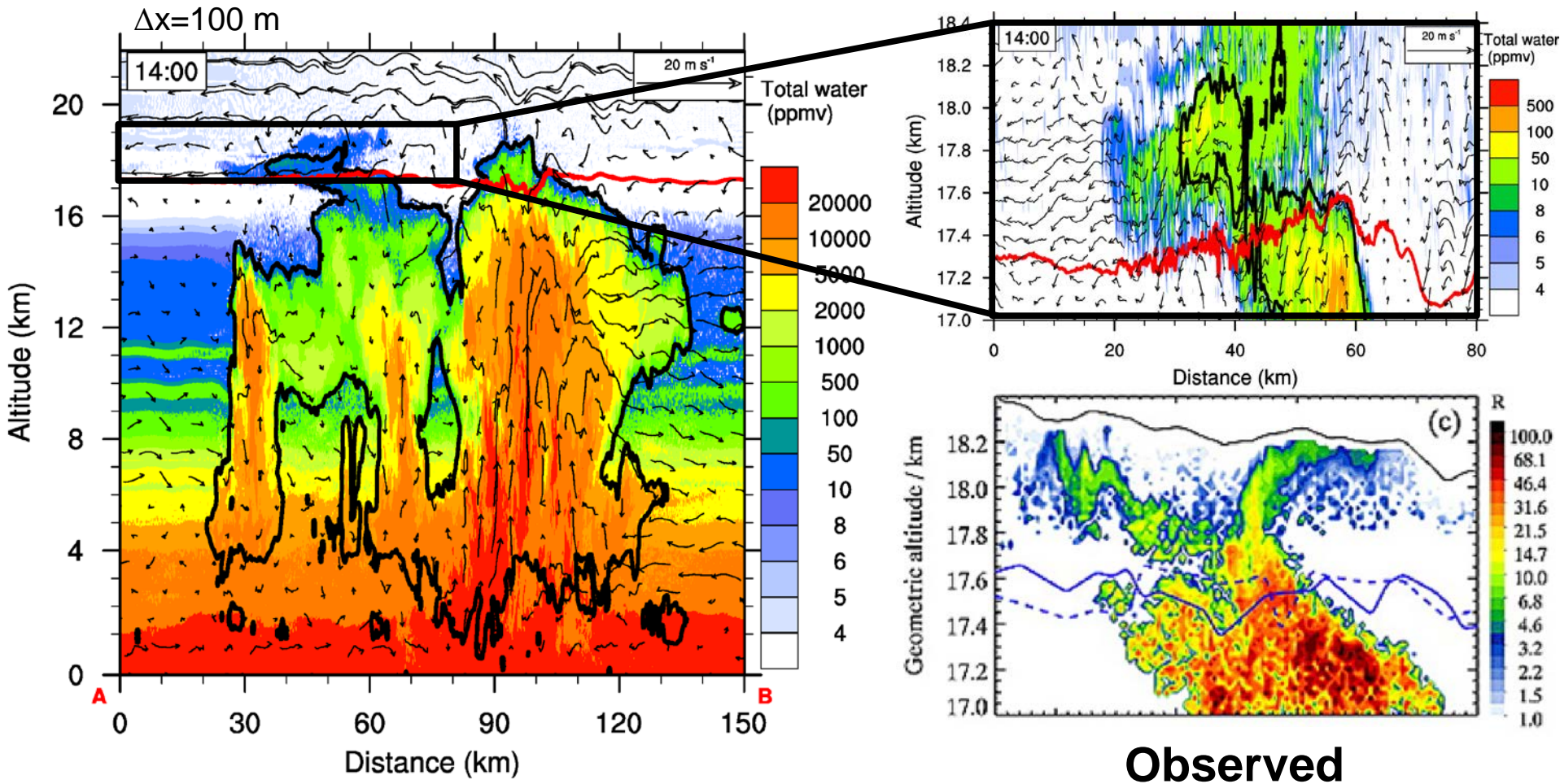


How does the very deep convection hydrate the stratosphere?

Video on <https://youtu.be/xjPumywGaAU>

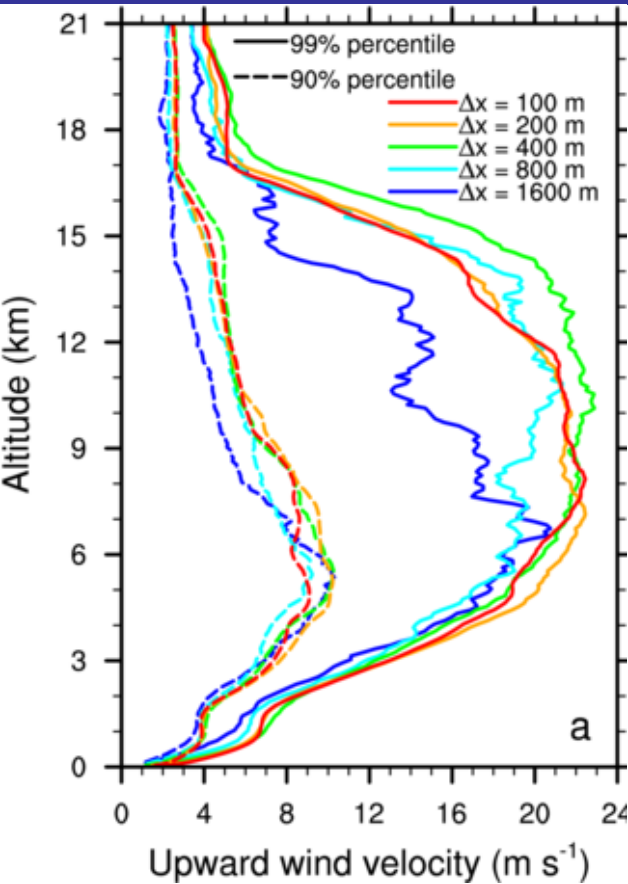
Dauhut et al., Atmos. Sci. Lett. 2015

Comparison with SCOUT-O3 observations

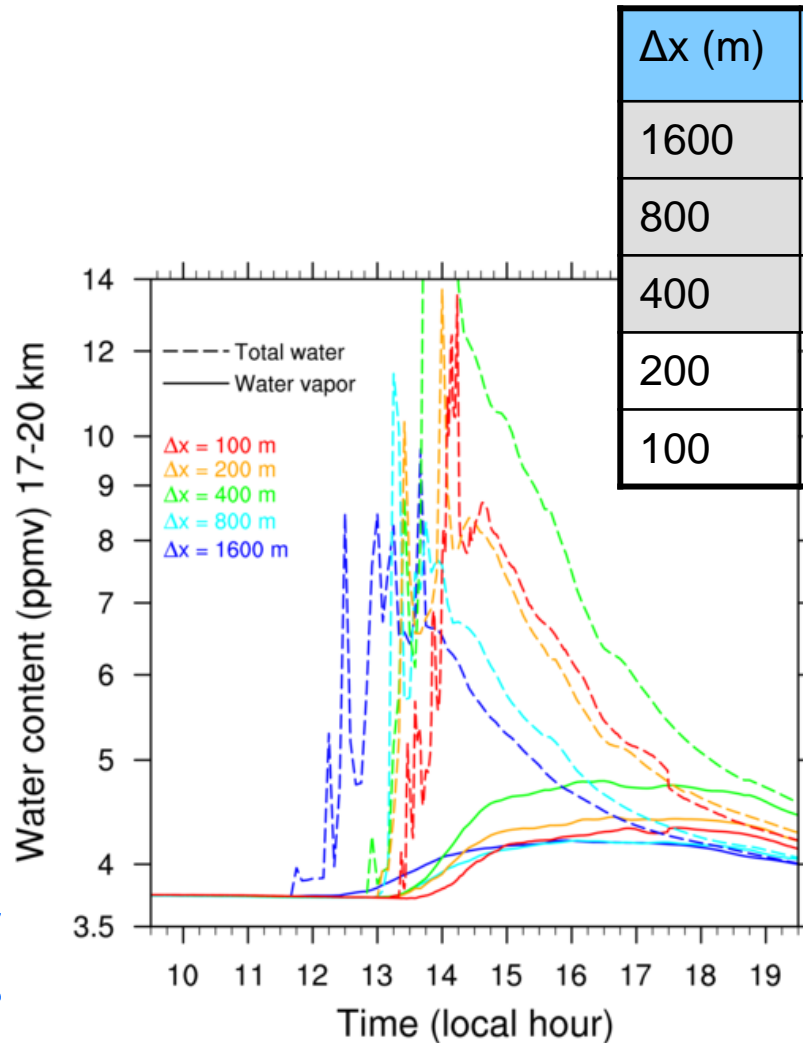


The Giga-LES reproduces correctly the details of Hector as its overshoots into the stratosphere

Sensitivity to horizontal resolution



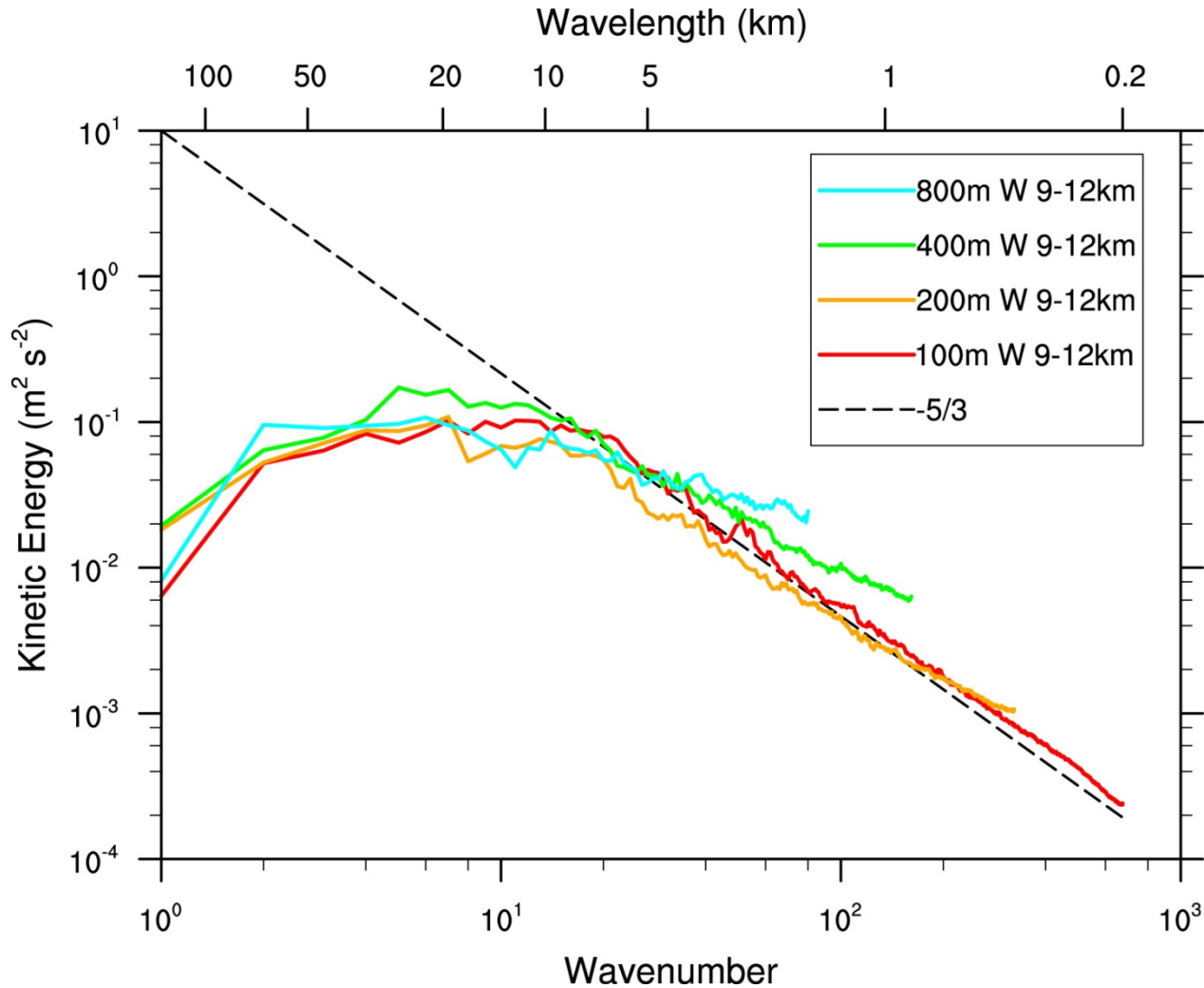
The vertical velocity for the most rapid updrafts generally decreases with reduced resolution



The magnitude of the moistening tends to converge with $\Delta x \geq 200$ m

Δx (m)	ΔH_2O (%)	ΔH_2O (t)
1600	+12	+2222
800	+13	+2252
400	+27	+4525
200	+18	+3263
100	+16	+2776

Spectrum of vertical velocity

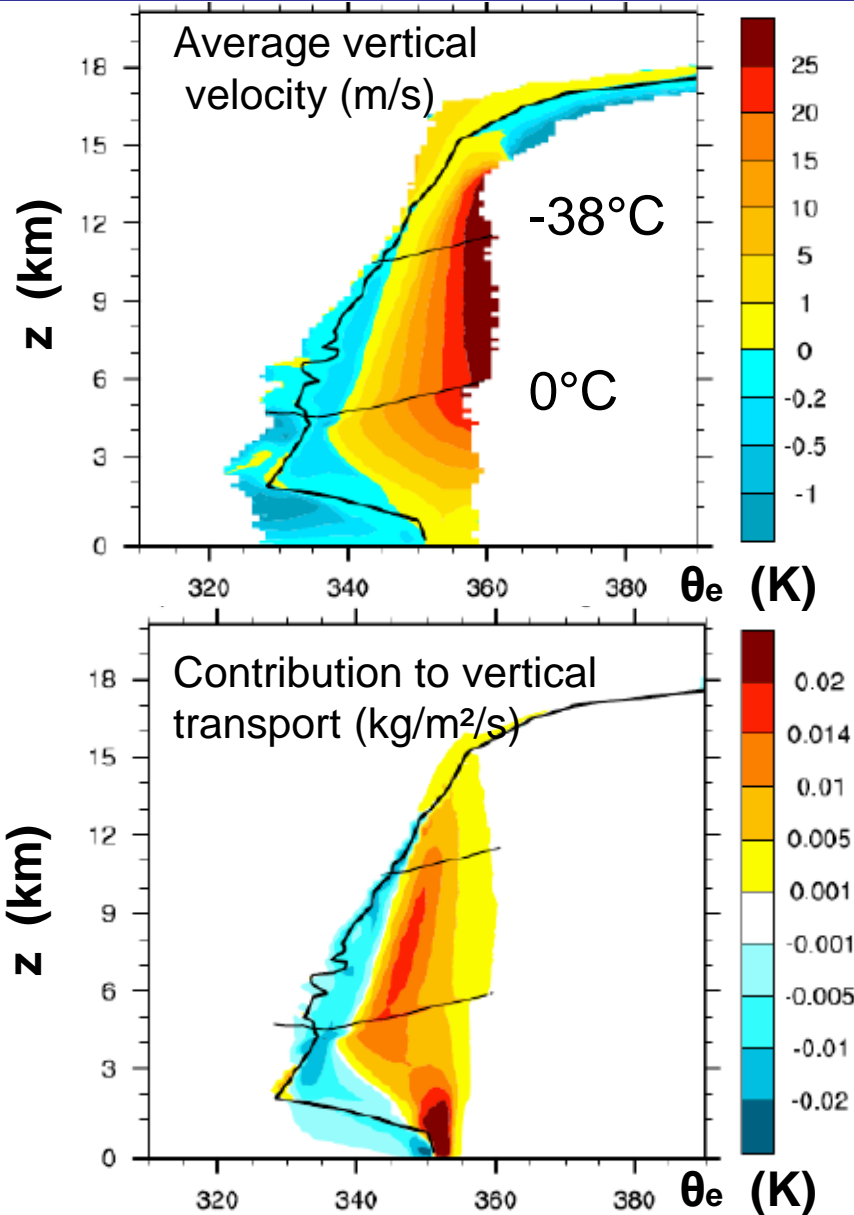


13:00

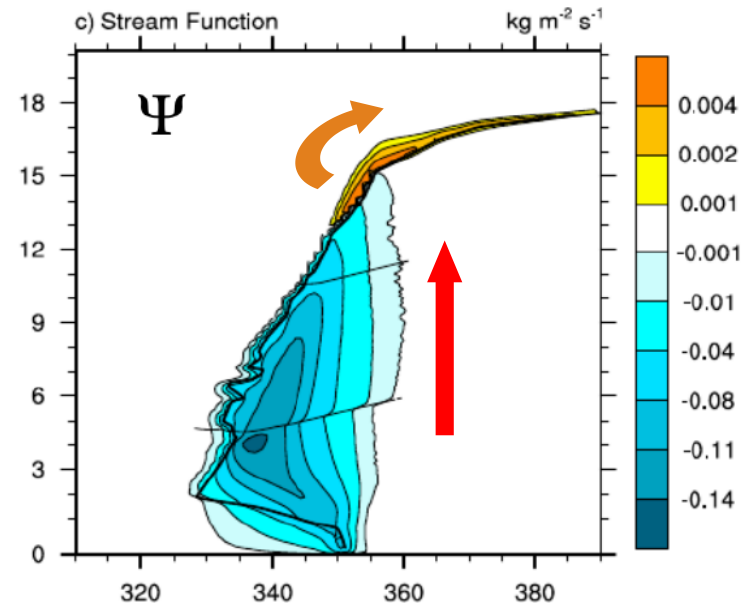
A grid spacing of $\Delta x=200$ m or 100 m is required for a reliable estimate of the hydration of the stratosphere

Overturning in Hector

Very deep convection
13:31 – 13:45



Two key circulations

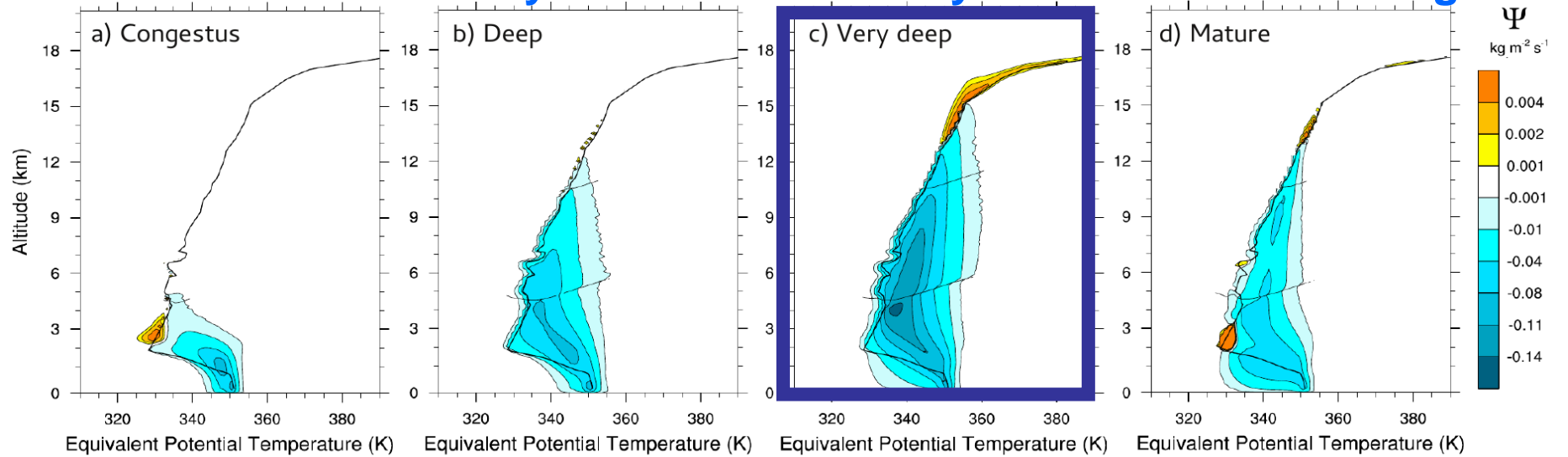


Overshoot overturning

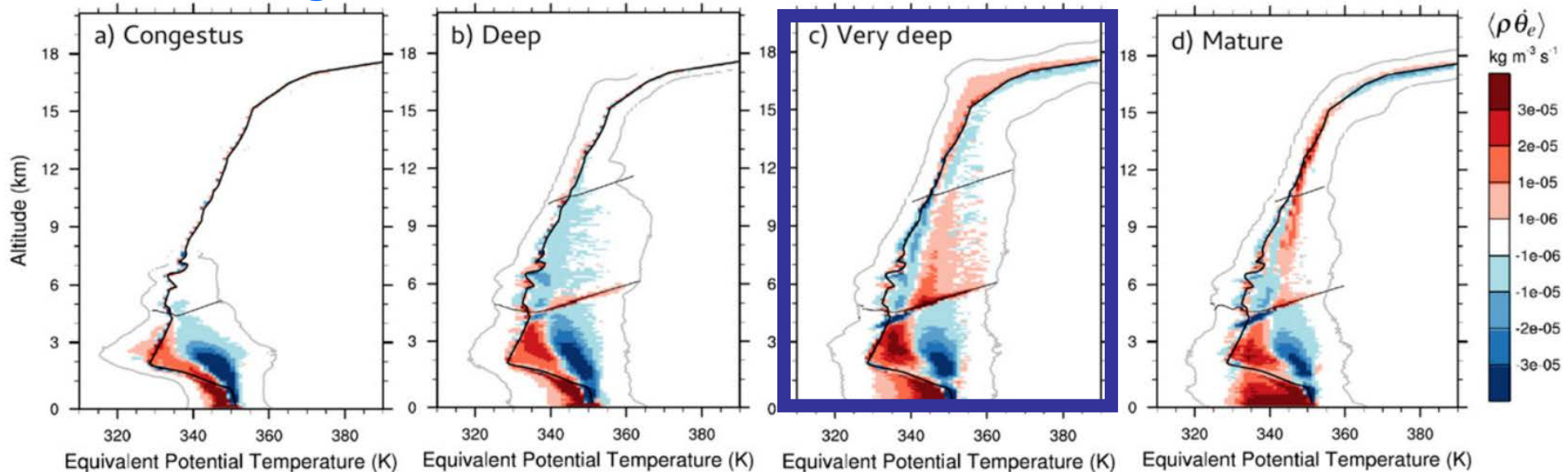
Tropospheric overturning

The very deep convective phase is

... characterized by a maximal intensity of the two overturnings

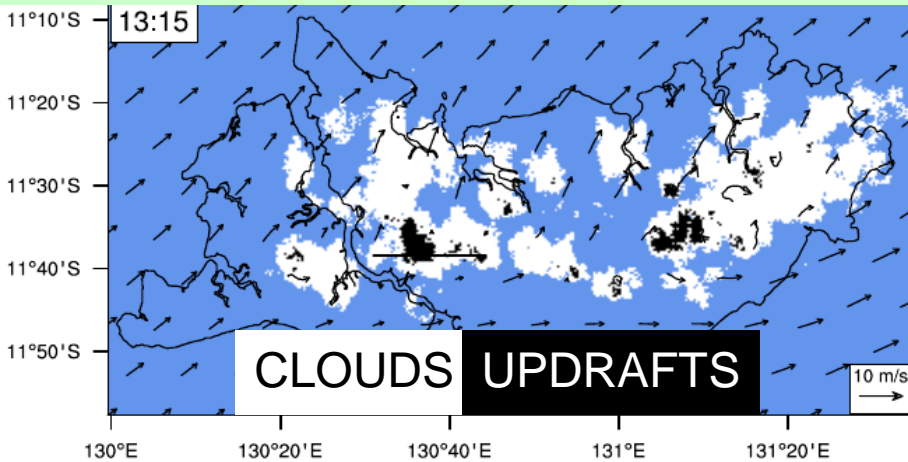


and large latent heat release due to ice formation

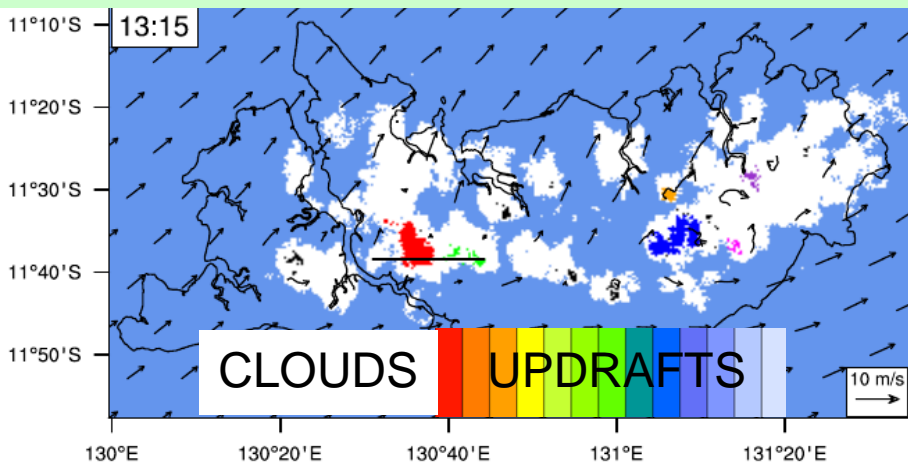


Identification of the tallest updrafts

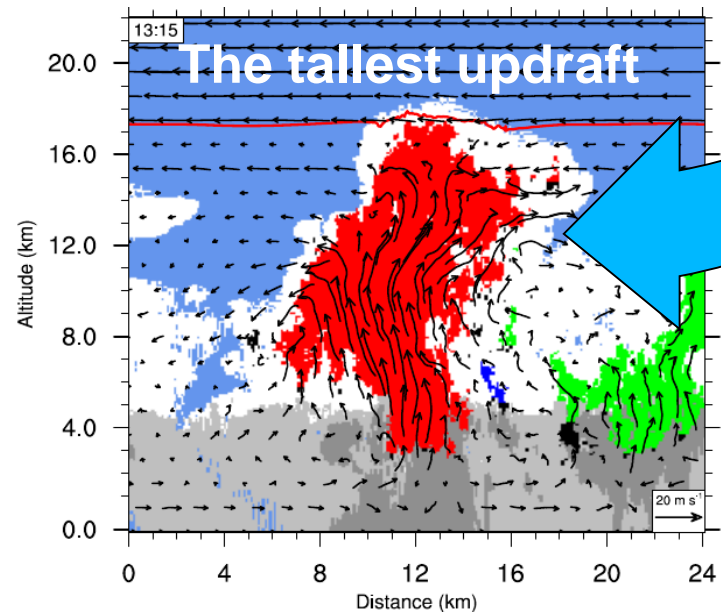
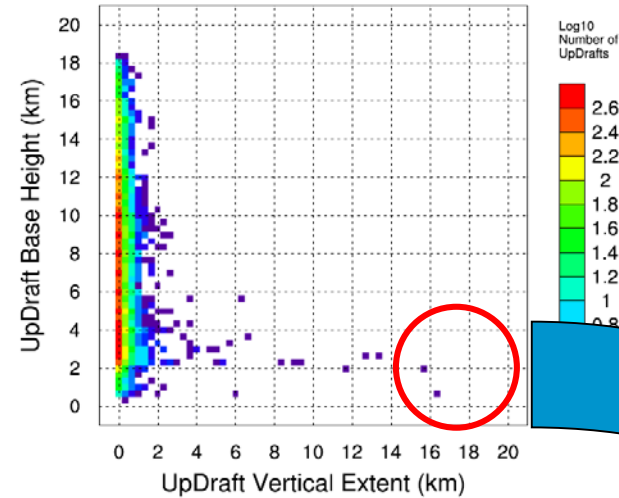
Step 1: detection of updrafts on every gridpoint where $w > 10$ m/s



Step 2: identification of updrafts as object

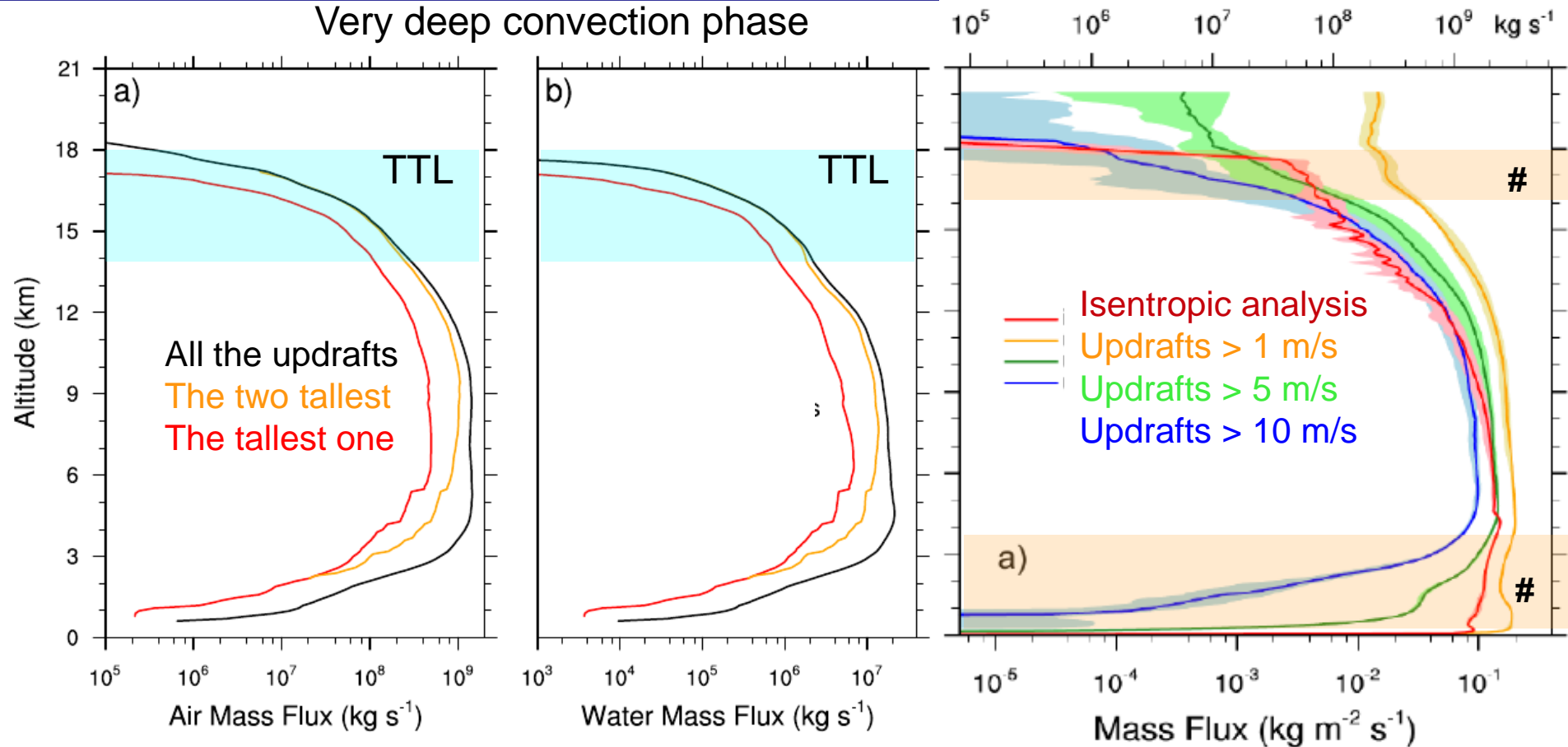


Step 3: Statistics of updrafts



The tallest updrafts, why bother?

Very deep convection phase



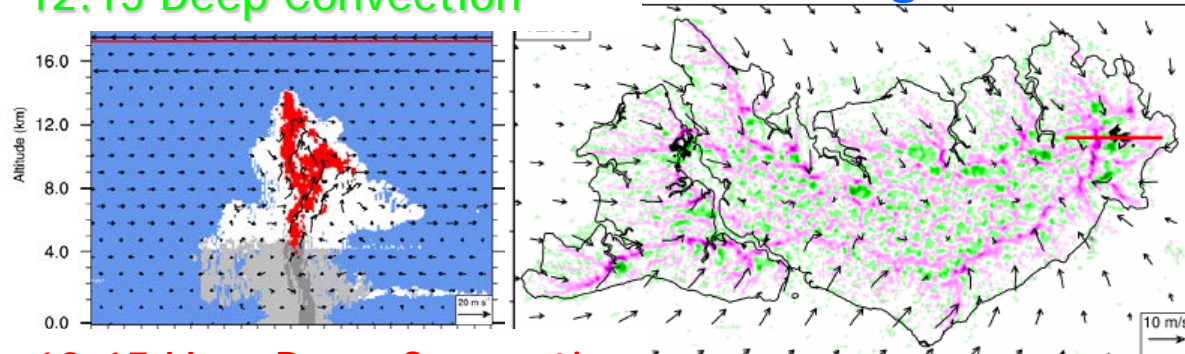
In TTL, the two tallest updrafts contribute to >90% of the transport by all the updrafts.

The isentropic analysis corroborates the Eulerian computation with $w > 10$ m/s, except in lower tropo and around the tropopause (#) where weak motions matter for the irreversible flux.

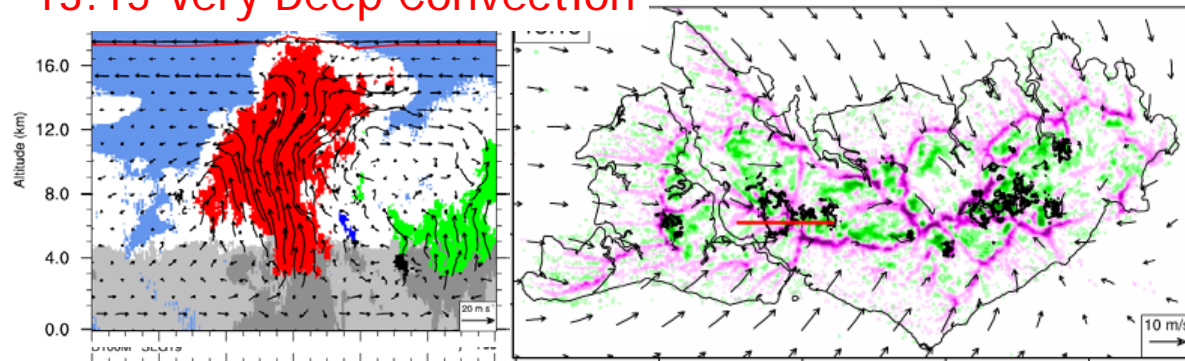
Formation of the tallest updrafts

Convergence intensified by cold pools

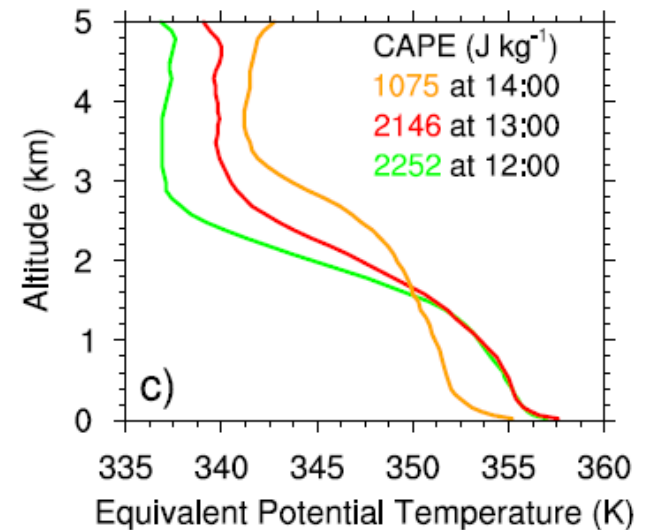
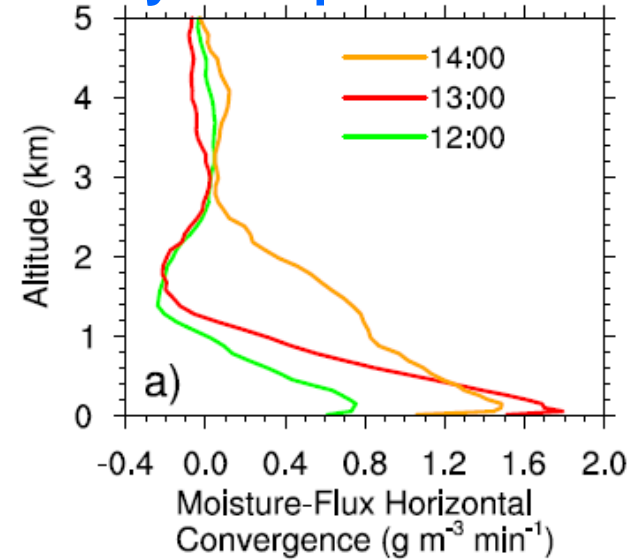
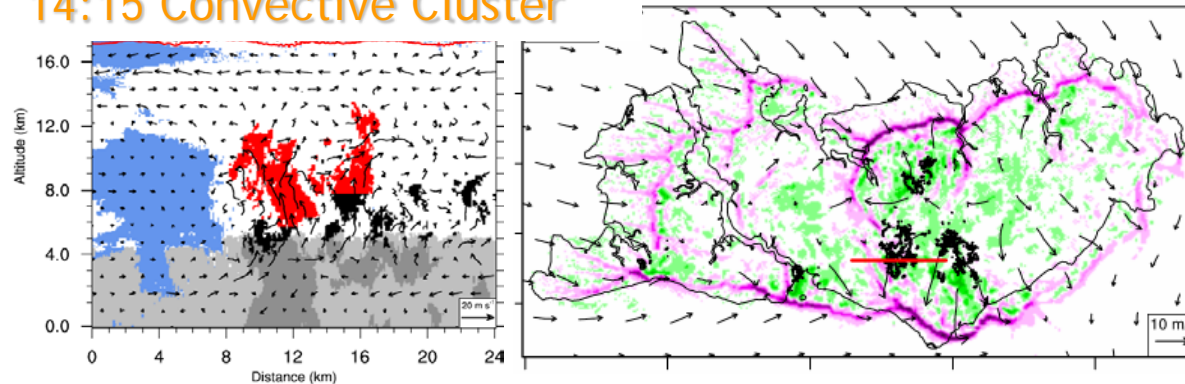
12:15 Deep Convection



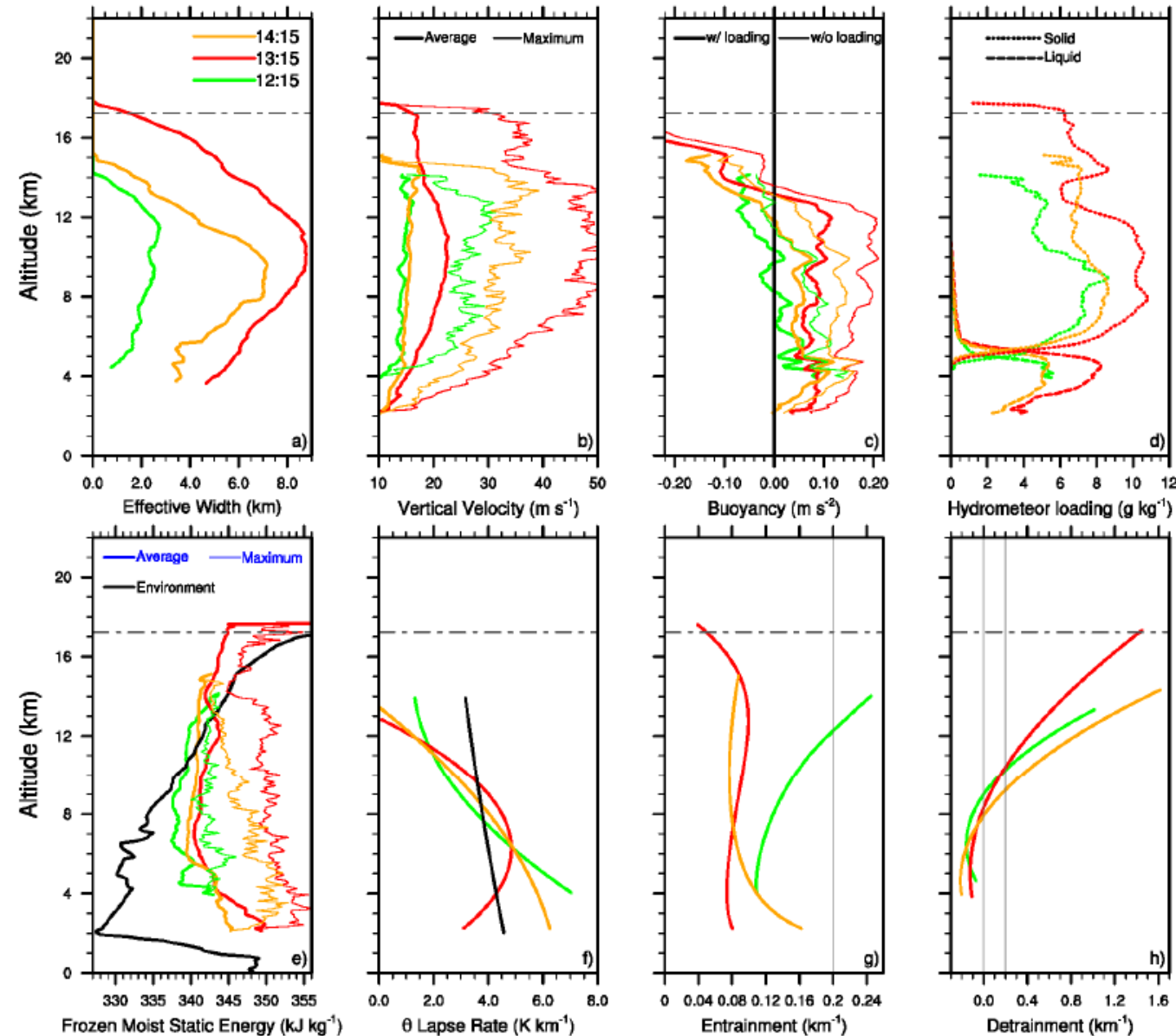
13:15 Very Deep Convection



14:15 Convective Cluster



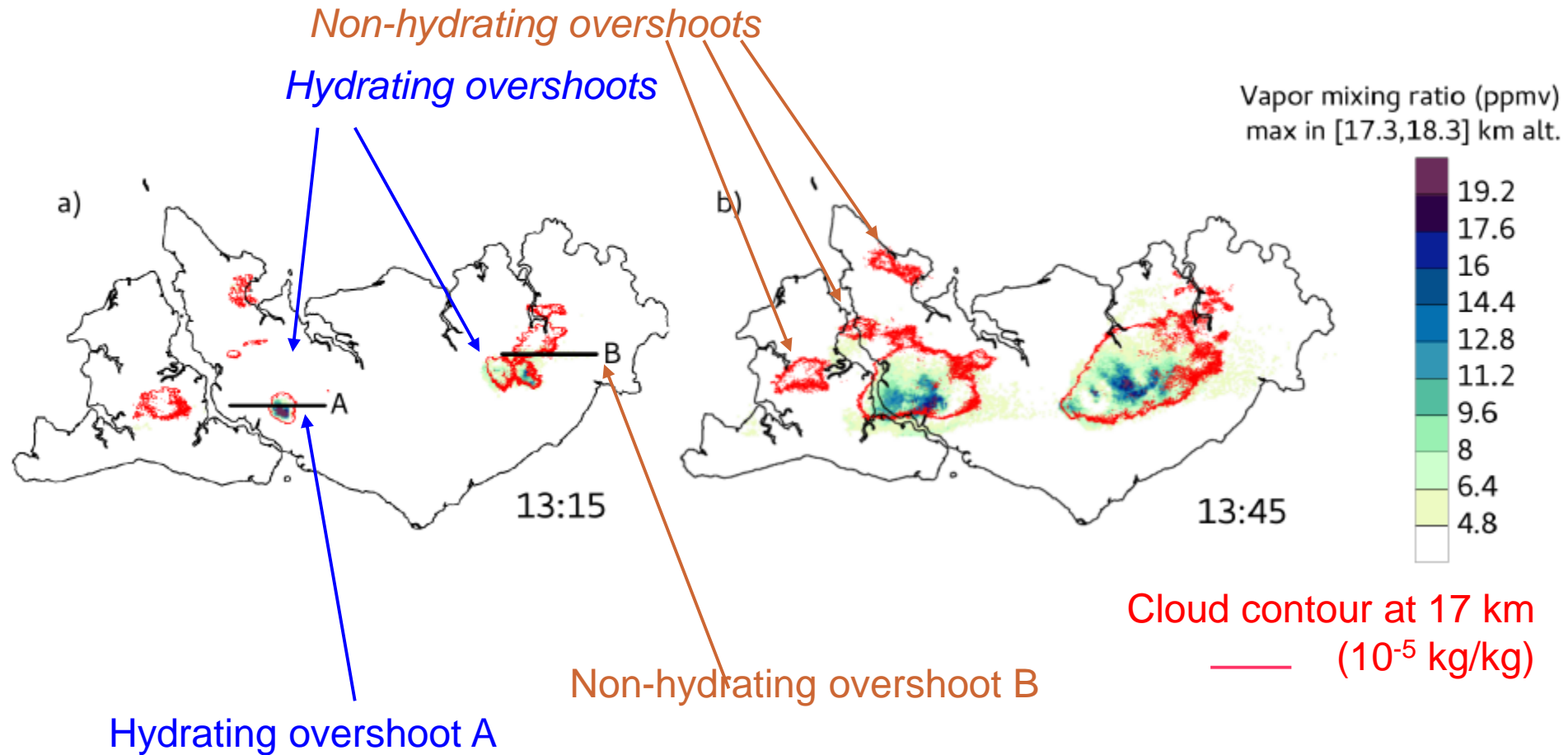
Properties of the tallest updrafts



The tallest updrafts that overshoot the stratosphere are larger, stronger, more buoyant, carrying more water, having more MSE, a larger lapse rate, less diluted than those occurred one hour earlier and after.

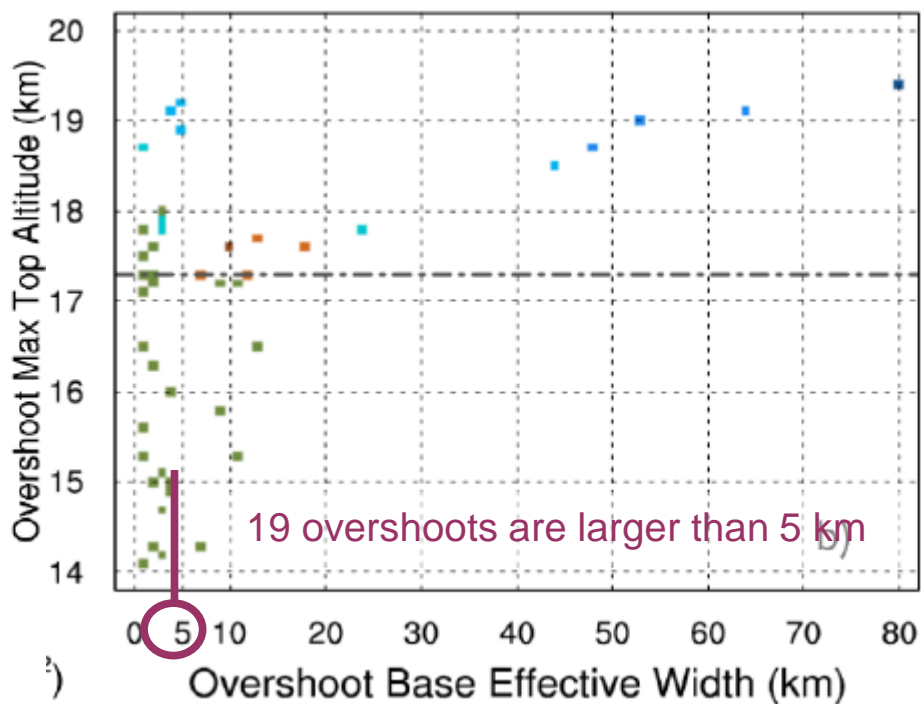
How do overshoots hydrate the stratosphere?

700 overshoots identified, 46 overshoots last more than 10 min
They are diverse in shapes and impacts above the tropopause



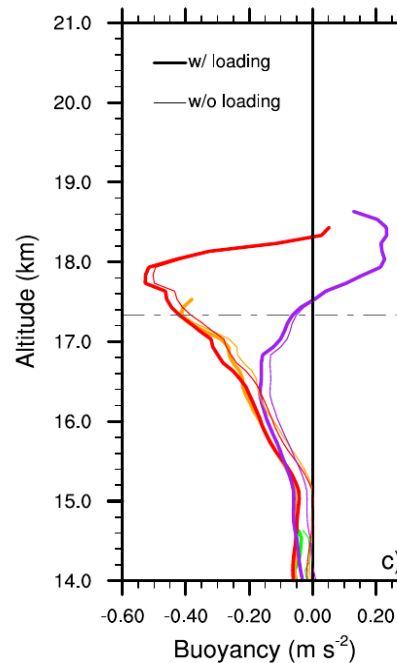
Hydrating and non-hydrating overshoots

The 46 overshoots that last > 10 min

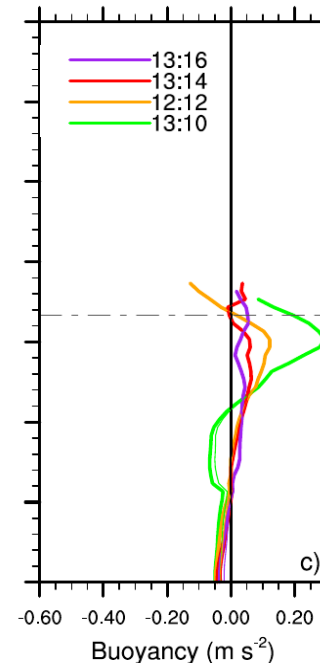


The hydrating overshoots reach higher altitudes...

Hydrating overshoot A

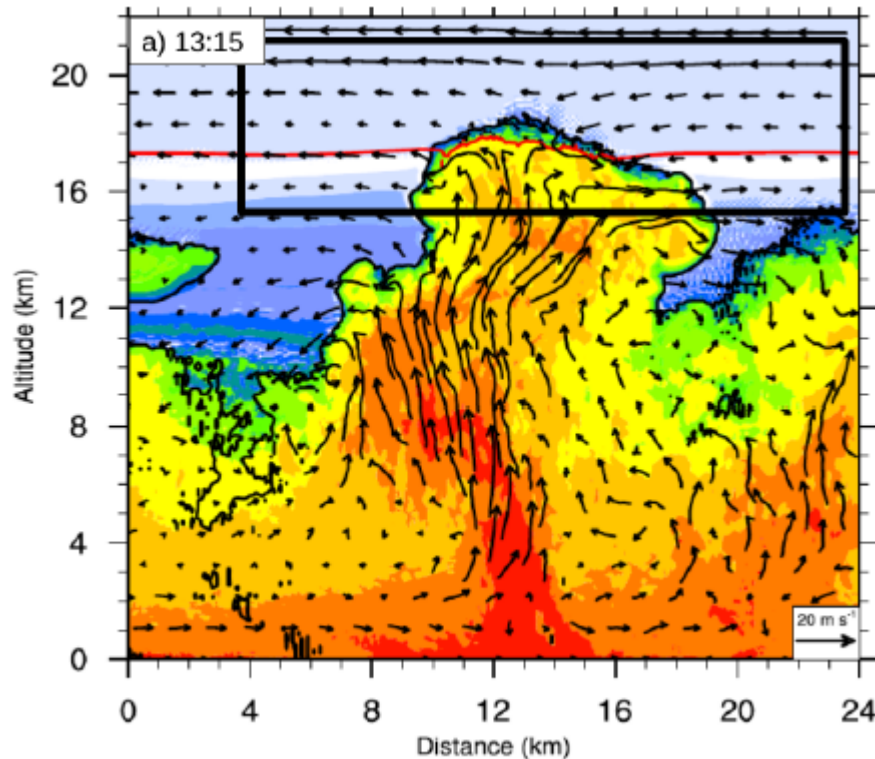


Non-hydrating overshoot B

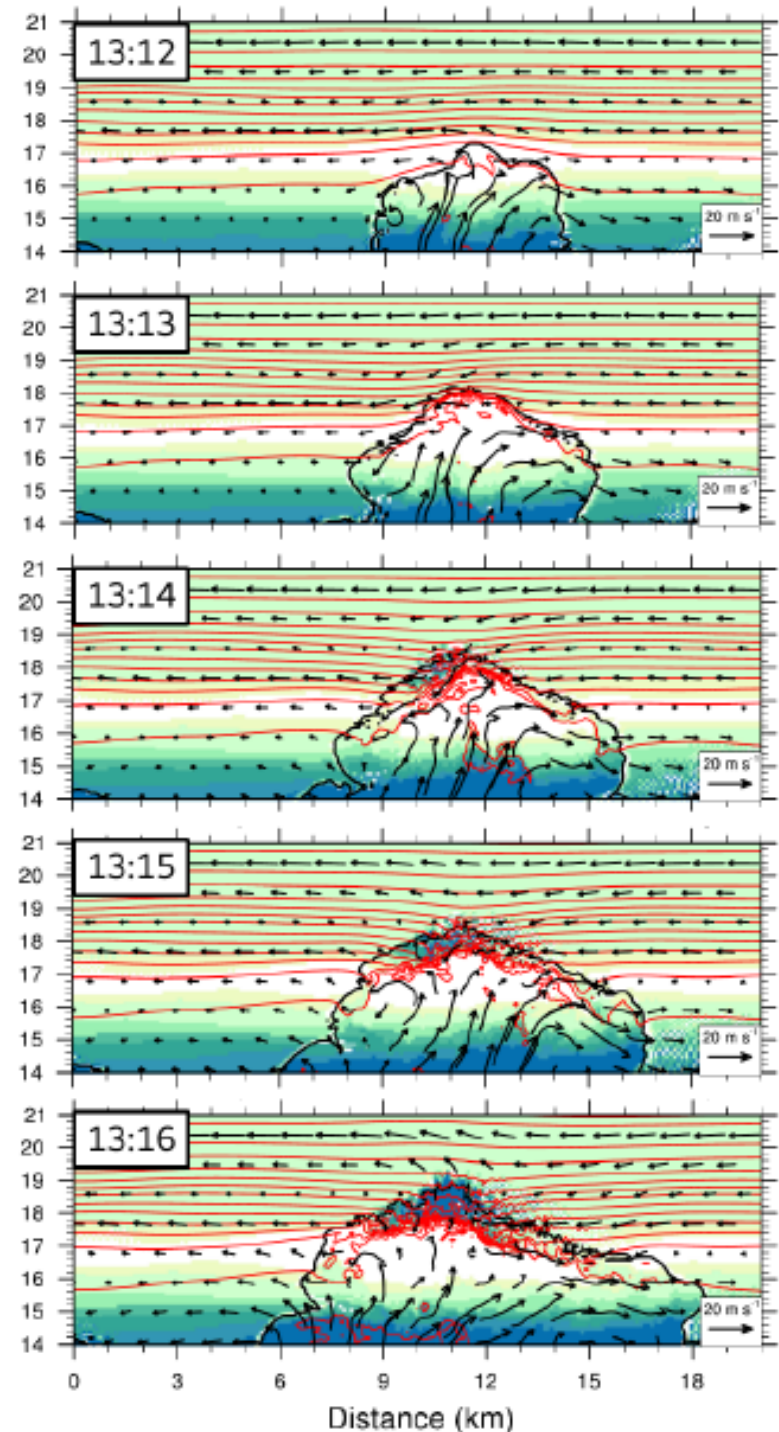
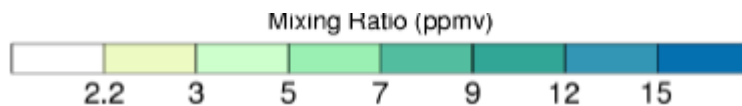


...and exhibit large absolute values of buoyancy.

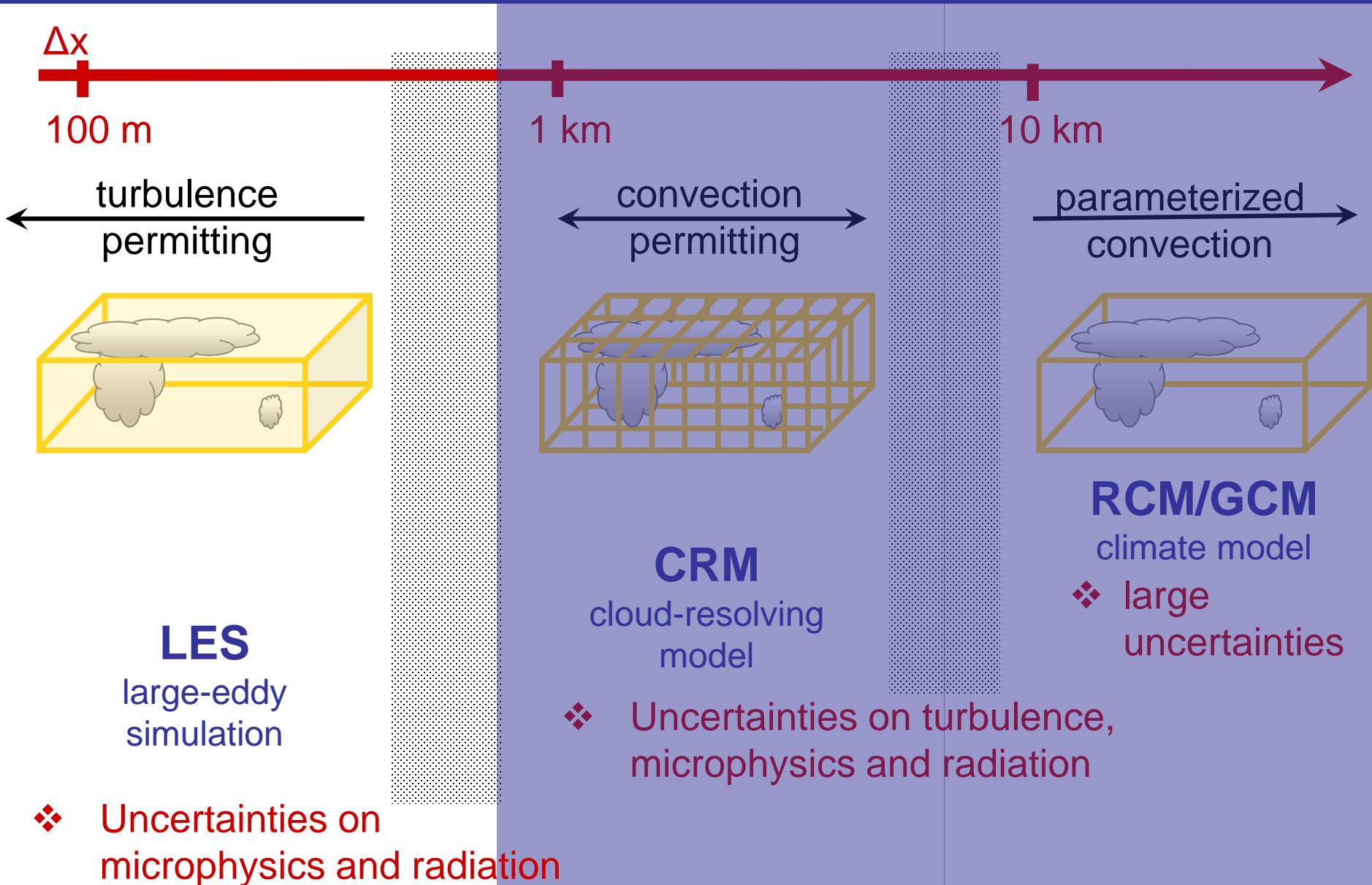
The mechanisms inside the hydrating overshoot A



The **top-entrainment** of stratospheric air is crucial to warm the overshoot and to produce the hydration by ice sublimation.

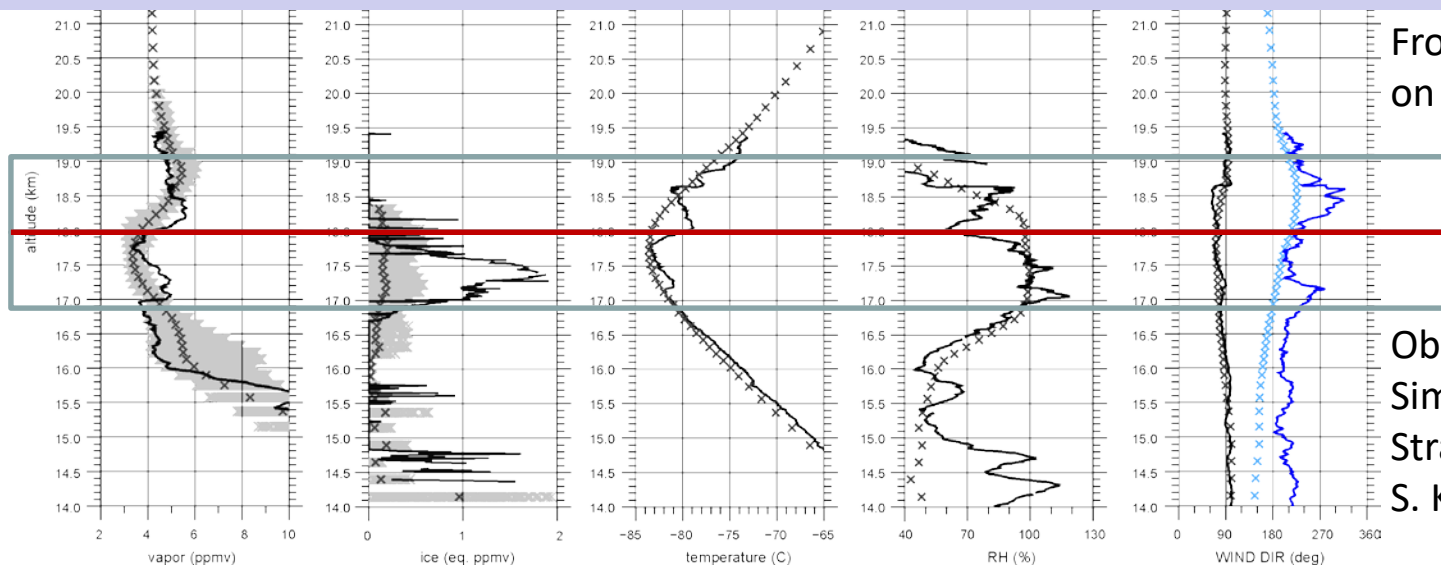


Representation of deep convection



Convective hydration during StratoClim

What is the fate of the water injected by overshoots?

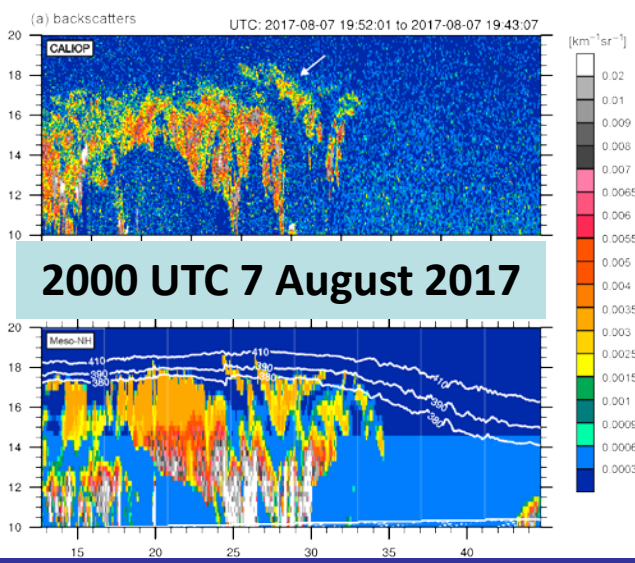
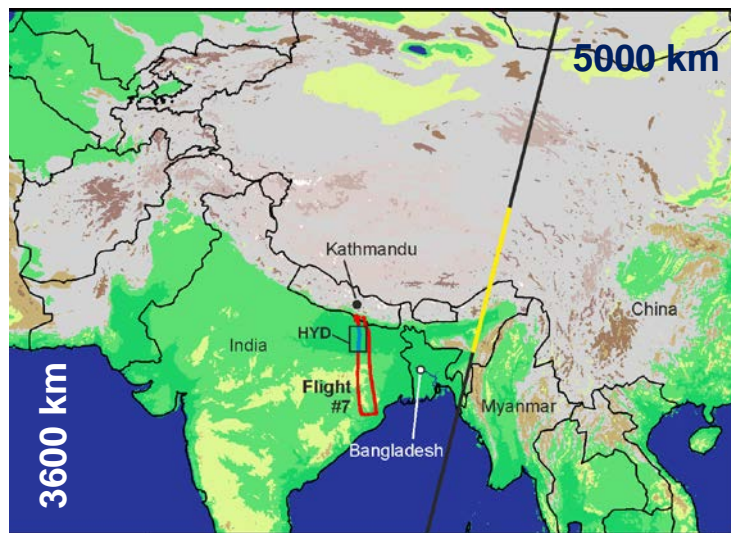


From 06:20 to 06:48 UTC on 8 August 2017

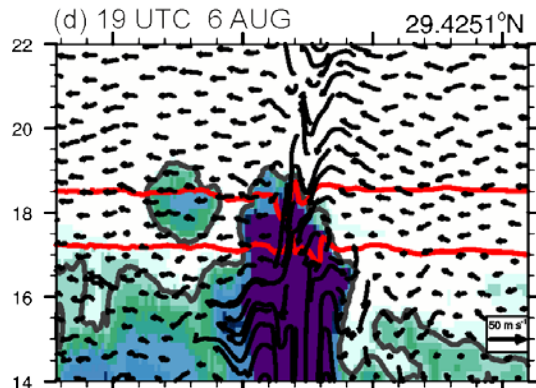
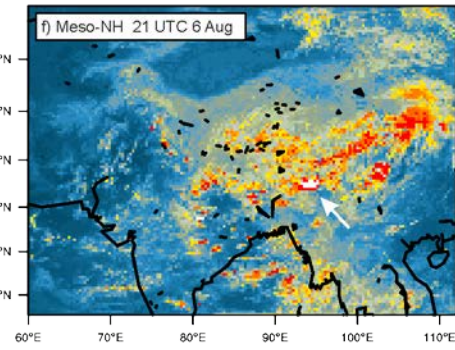
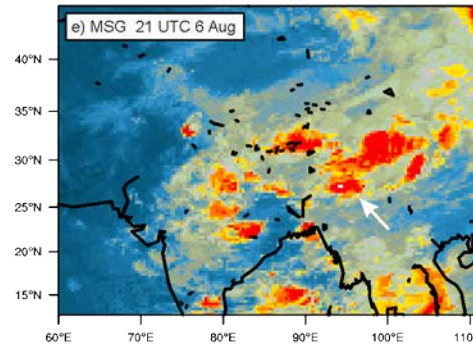
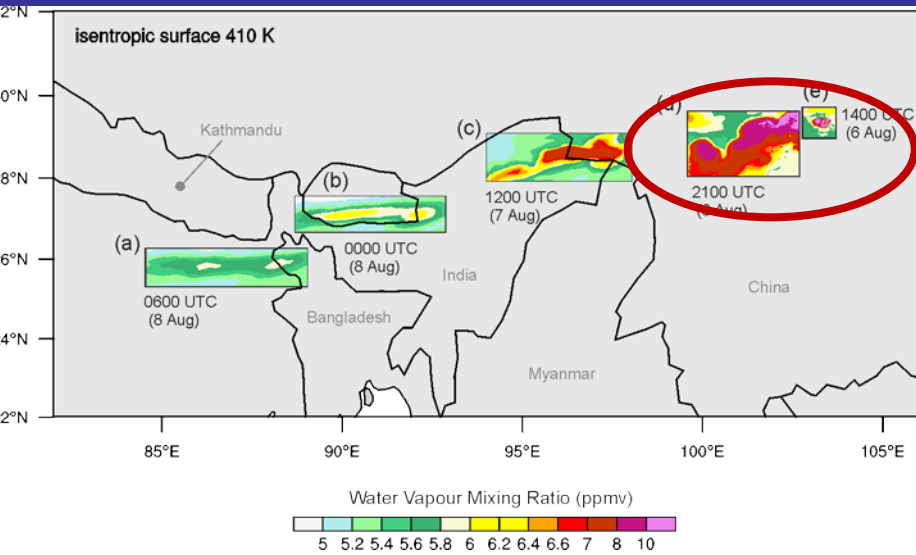
Tropopause

Observation: solid line
Simulation: cross marks
StratoClim data courtesy of S. Khaykin and M. Krämer

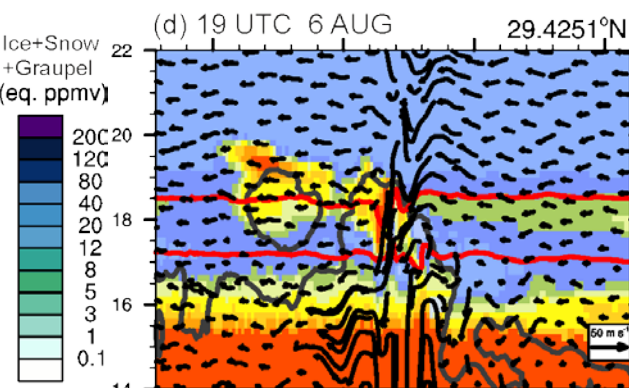
Simulation starting at 00 UTC 6 August 2017 from ECMWF analysis and run for 3 days, $\Delta x=2.5$ km, 2000 x 1440 x 144, 400 million gdpts



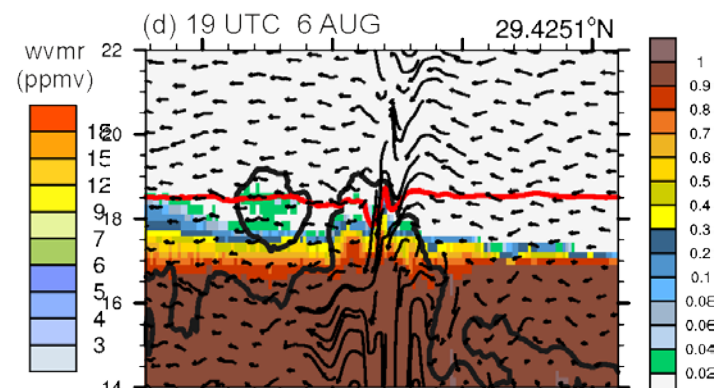
Injection of water by overshoots



Ice content

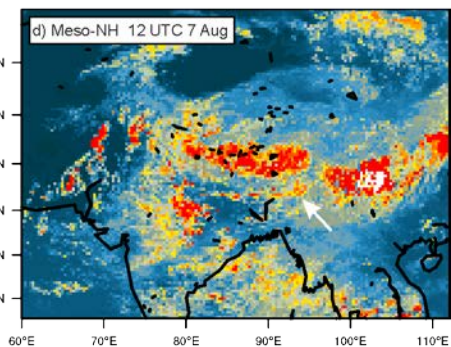
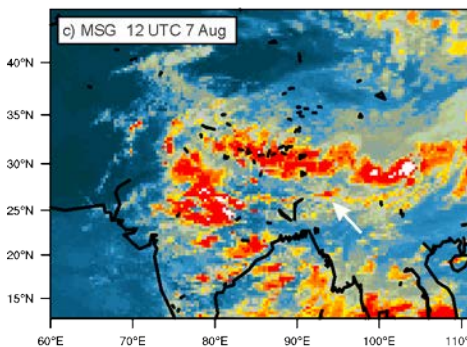
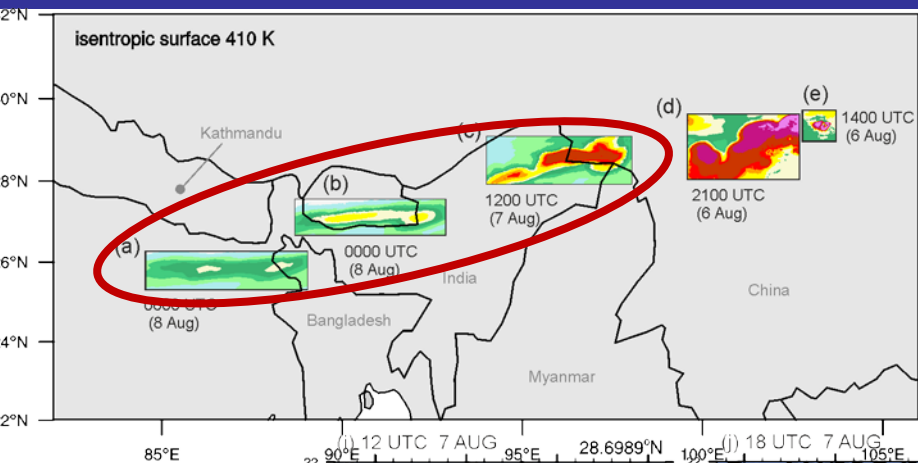


Water vapor

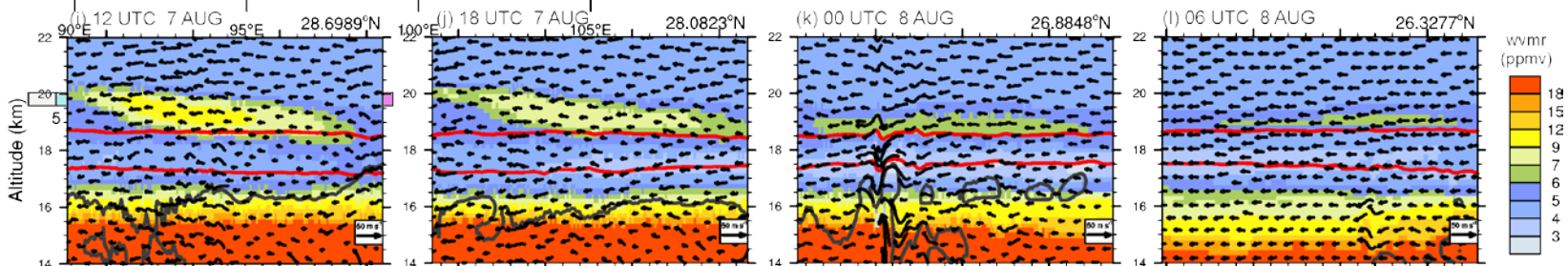


Tropospheric tracer

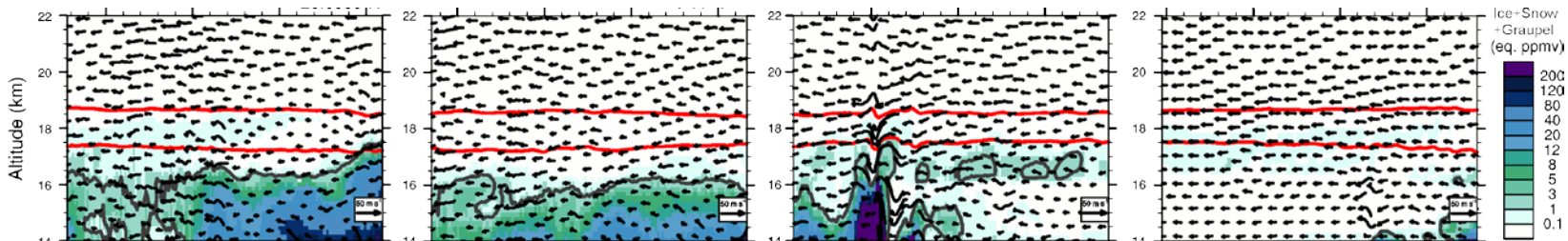
Advection of the hydration patch



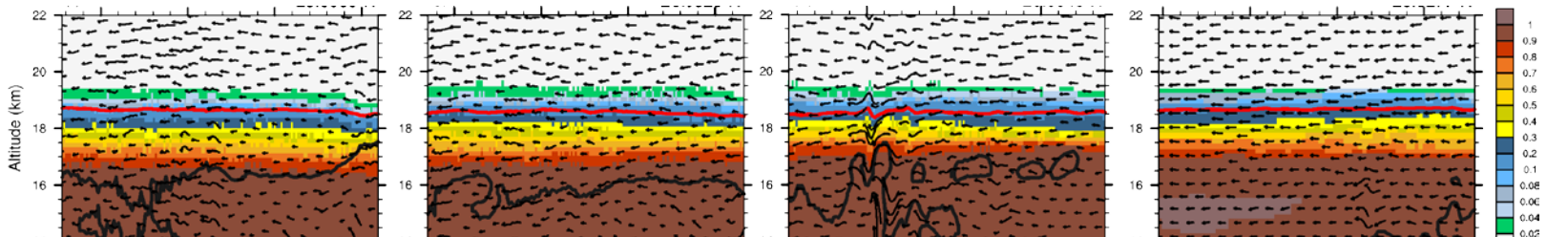
Water vapor



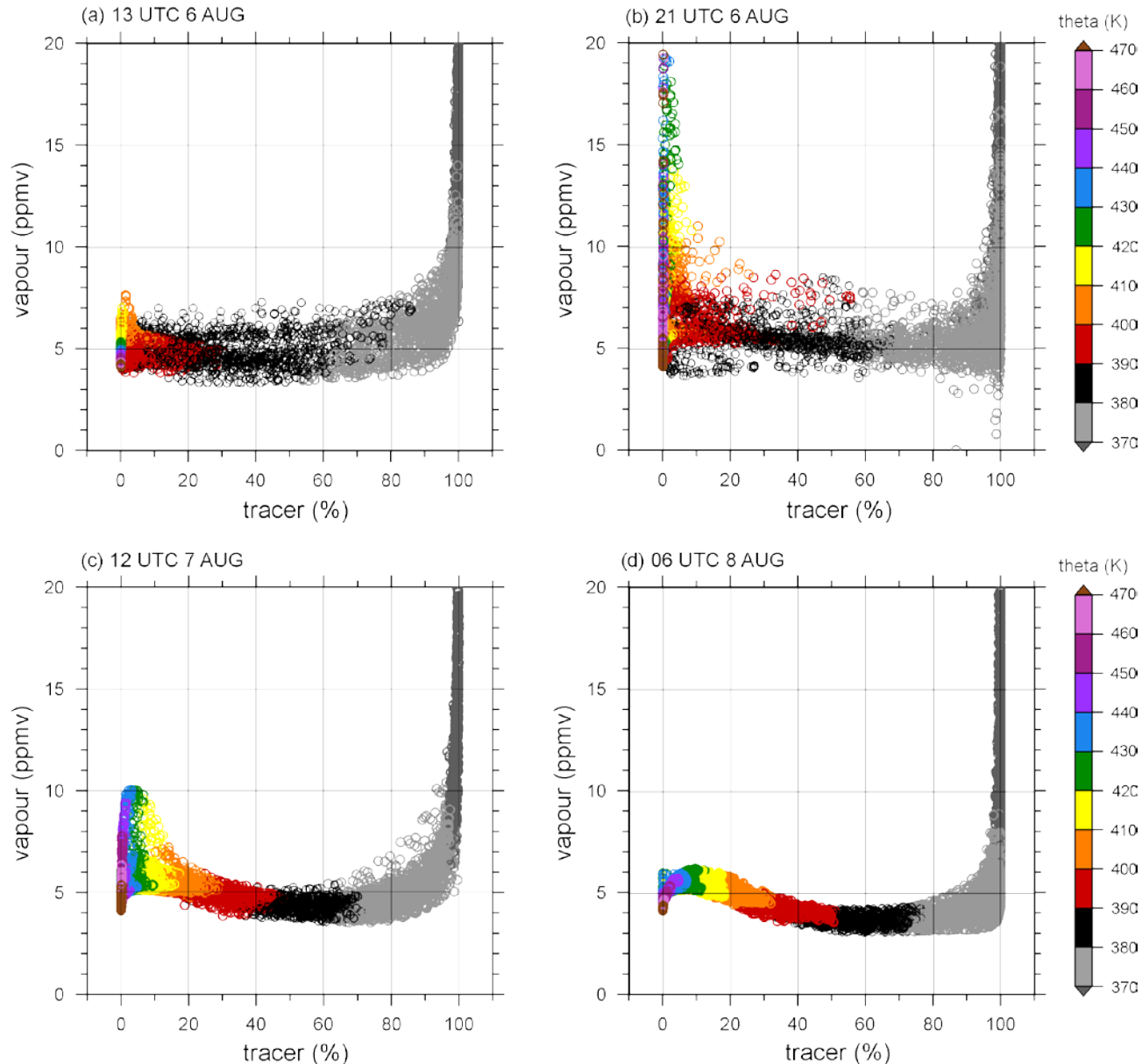
Ice content



Tropo. tracer

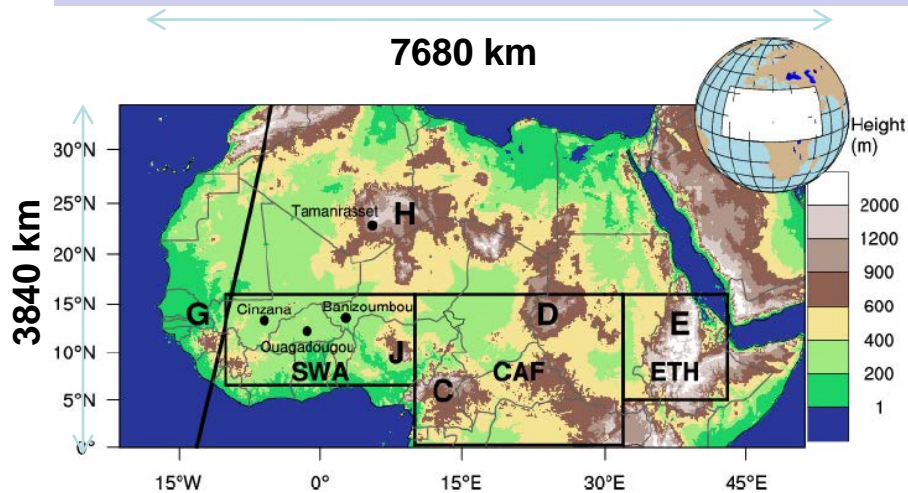


Convective vs. turbulent mixing

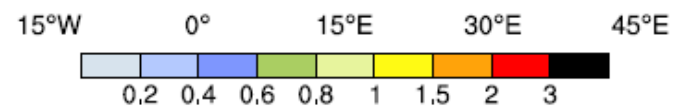
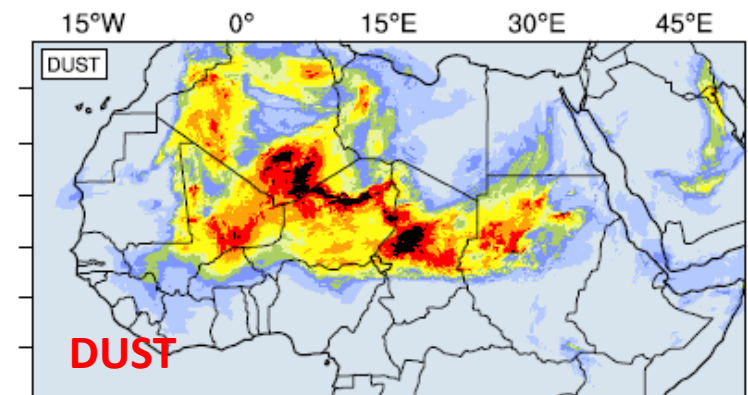
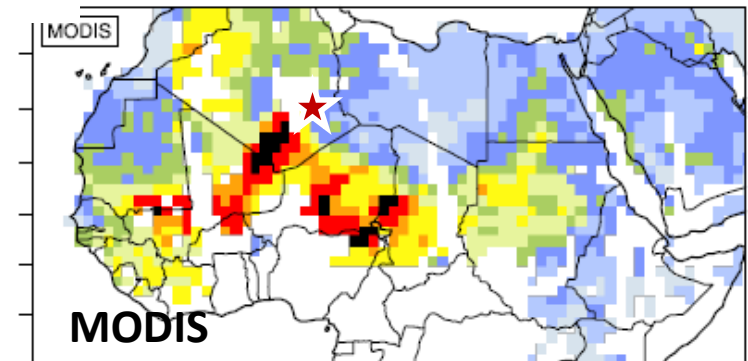


MCSs over north Africa

- *What controls the distribution and variability of precipitation?*
- *What is the radiative impact of dust on the atmosphere?*



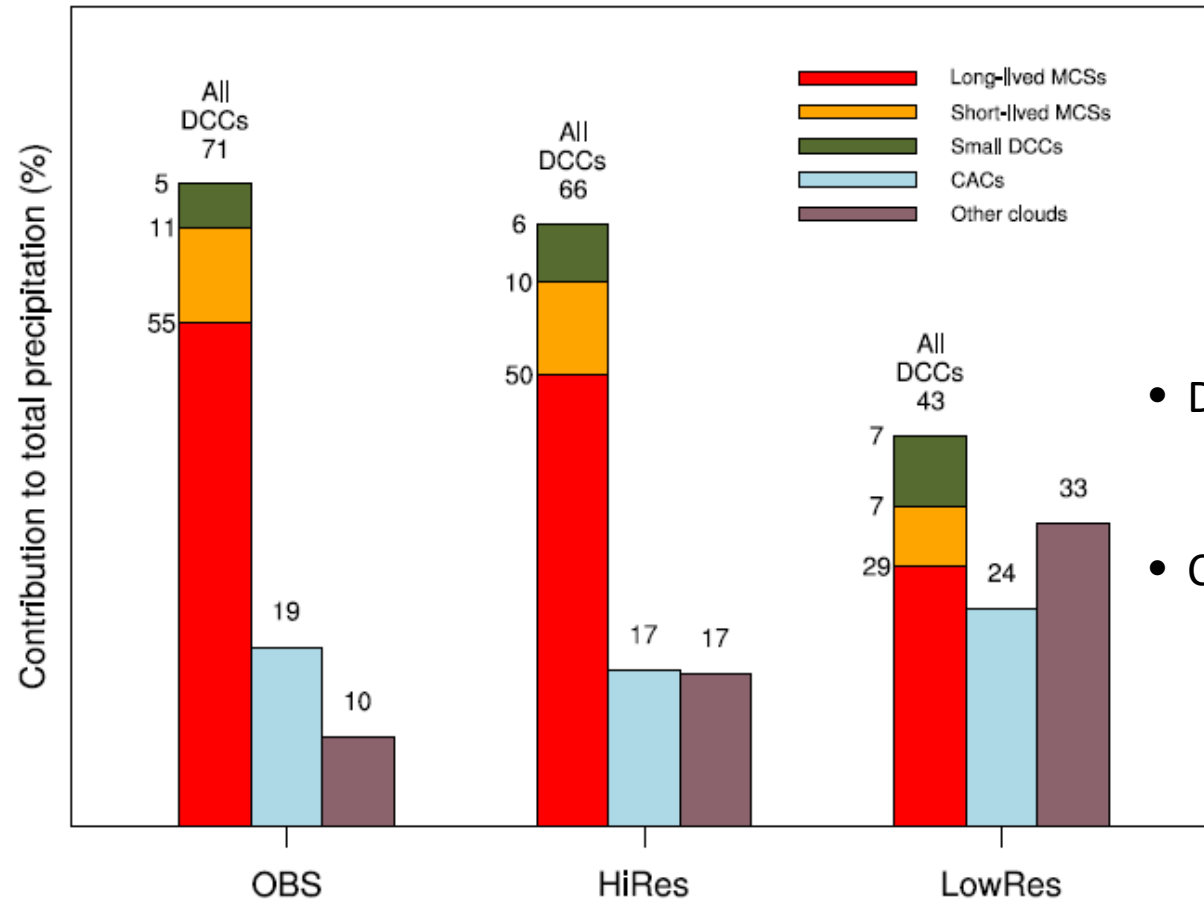
11 June 2006 around 1200 UTC



Three simulations starting at 00 UTC 9 June 2006 from ECMWF analysis and run for 6 days

- **HiRes/DUST** $\Delta x=2.5$ km, 3072 x 1536 x 72, 1/3 billion gdpts, with dust radiative effects
- **LowRes** $\Delta x=20$ km with KFB convective parameterization and dust scheme
- **NODUST** $\Delta x=2.5$ km, w/o dust radiative effect

Distribution of precipitation

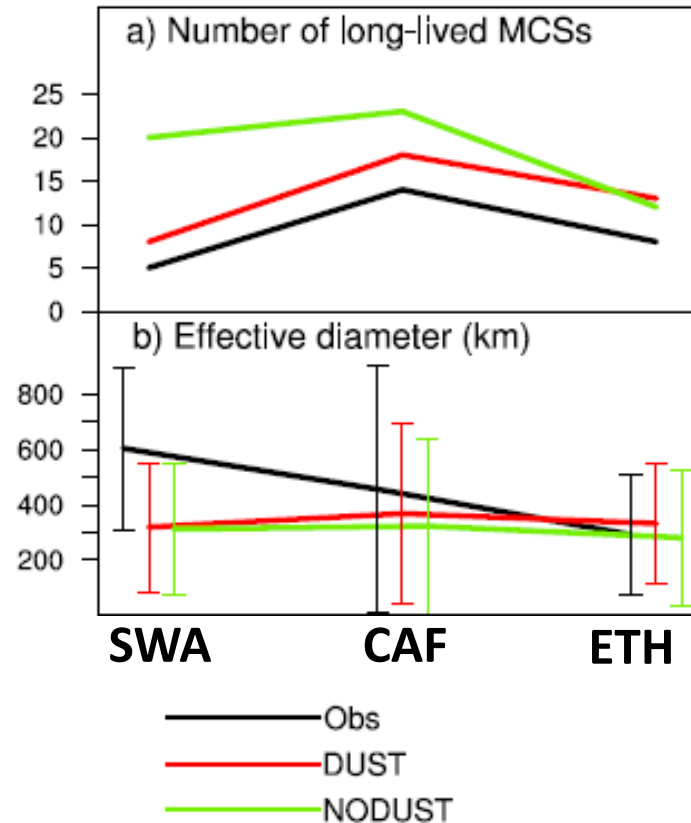
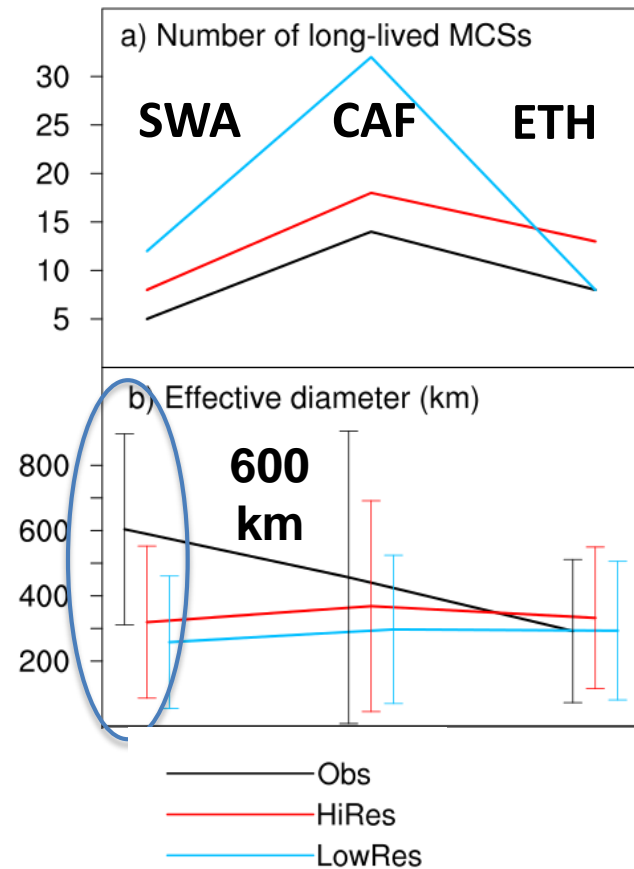


OBS: MSG 10.8 μm brightness temperature and TRMM 3B42 rain product

- Deep convective clouds: $\text{BT} < 230\text{K}$
- **MCSs** = large DCCs ($\text{Deff} > 120\text{km}$)
- **Long-lived MCSs:** duration $> 6\text{h}$
- Cirrus anvil clouds $230\text{K} < \text{BT} < 260\text{K}$

❖ Long-lived MCSs produce 55% of precipitation, thus a large part of the diurnal cycle. This is well captured by the CRM simulations

Characteristics of long-lived MCSs



OBS: most organized long-lived MCSs in SWA

HiRes/DUST agreement with OBS except in SWA (too small, short-lived and slow, small northward meridional component)

LowRes drawbacks more pronounced

DUST long-lived MCSs less numerous than **NODUST** in SWA

- ❖ The CRM simulations lack the degree of organization of the long-lived MCSs over SWA. The CRM with dust does a better job due to the stabilization of the lower atmosphere (large CAPE), which inhibits the triggering of convection (large CIN).

Conclusions

- ✓ CRM approach was successful in representing cloud and precipitation distribution – MCSs, diurnal cycle, etc.
- ✓ A higher skill was obtained with dust radiative effects over north Africa, but cloud organization is lacking: errors in initial conditions, drawback in parameterization of microphysics, of turbulence?
- ✓ Convergence in dynamics, hydration is almost reached with LES

❖ Future plans

- ✓ Case studies of deep convection using an aerosol-aware microphysical scheme – avoiding the saturation adjustment and looking for aerosol-cloud interactions: Application to the AEROCLO-sA/ORACLES/CLARIFY field campaign (Aug. 2017)
- ✓ Investigation of convective overshoots over longer periods