



Strategic Implementation Plan (SIP) for a Community-based Unified Modeling System



Dynamics and Nesting WG

Presented by

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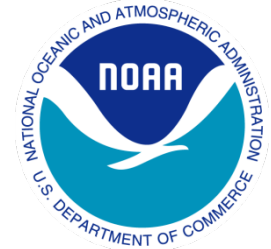
Presented at NOAA Community Modeling Workshop

April 18-19, 2017; College Park, MD



Dynamics & Nesting WG

Membership



SIP WG Team Members

Last Name	First Name	Org
Tallapragada	Vijay**	NCEP/EMC
Harris	Lucas**	GFDL
Gopalakrishnan	Sundaraman**	HRD/AOML
Jablonowski	Christiane**	U. of Michigan
Black	Tom	NCEP/EMC
Trahan	Samuel	NCEP/EMC
Mehra	Avichal	NCEP/EMC
Juang	Henry	NCEP/EMC
Viereck	Rodney	NCEP/SWPC
Yudin	Valery	CIRES/CSU
Mahajan	Rahul	NCEP/EMC
Kleist	Daryl	NCEP/EMC
Fuller-Rowell	Tim	NCEP/SWPC

NGGPS D&N Team Members

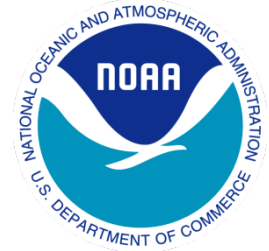
Last Name	First Name	Org
Wicker	Lou	NSSL
Sun	Shan	ESRL/GSD
Govett	Mark	ESRL/GSD
Putnam	Bill	NASA/GMAO
Goldhaber	Steve	NCAR/CGD/CESM
Zhang	Xuejin	HRD/AOML
Liu	Fei	NESII/ NEMS
Lin	Shian-Jiann	GFDL
Reinecki	Alex	NRL Monterey
Wang	Ning	ESRL/GSD
Jovic	Dusan	NCEP/EMC
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Bender	Morris	GFDL
Doyle	Jim	NRL Monterey

***SIP D&N WG Co-Chairs*

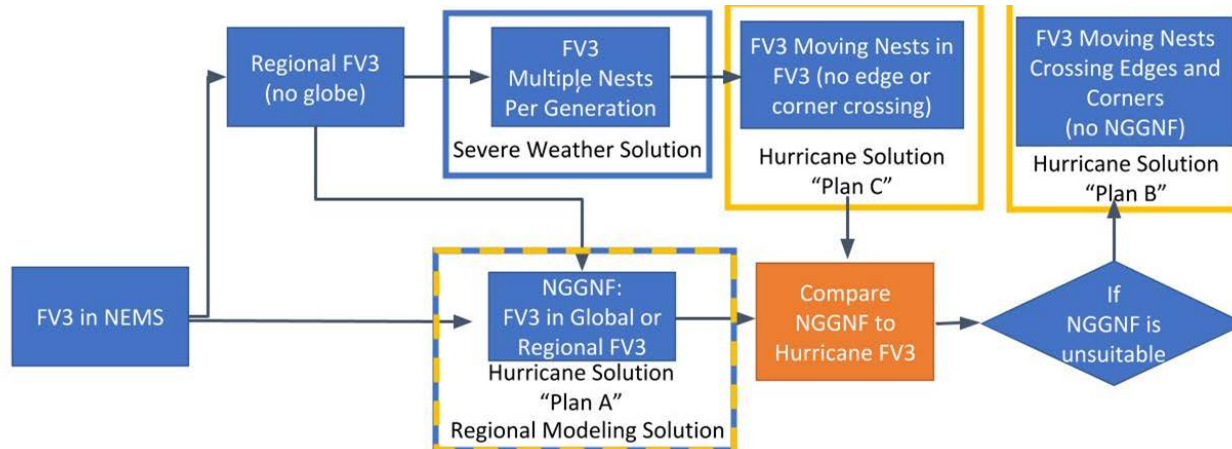


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Plans for Static and Moving Nests in FV3



- Path forward is clear. This is a “doing” group not a “talking” group. Tasks include:
 - *Implementation of moving nests for hurricane modeling purposes. Two potential solutions have been identified and will begin development.*
 - *Further improvements to nesting and stretching for regional applications, especially for high-resolution CAM modeling and WoF.*
 - *Model infrastructure (NEMS and FMS) support for these tasks*
 - *Computational efficiency and ease of use for research and operations drive the developments*





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Plan for Deep Atmosphere Dynamics



Scope: A deep atmosphere and variable-composition option, consistent with existing FV³ algorithm, for space weather prediction.

Philosophy: Minimizing changes to FV³ dycore and introduce changes incrementally; utilize what we have done previously with WAM-GSM, except maintain use of potential virtual temperature rather than enthalpy

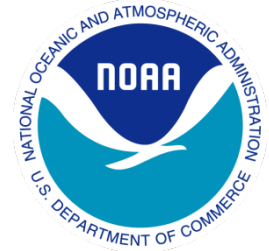
Requirements and Plan:

- Variable gas constant R_i & specific heats Cp_i with five tracers
- Extend vertical extent to $\sim 10^{-7}$ Pa (~ 600 km) and add additional physical processes, including strong 3D molecular diffusion, interaction with plasma (ionosphere) on different magnetic field grid system, regridding with ESMF
- Deep atmosphere dynamics, variable geocentric distance r , gravity $g(z)$, and extra terms in Coriolis force
- Ensure tolerance for $T > 2000$ K and $U, V > 1000$ m/s, $W > 100$ m/s



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Key issues to resolve



- How can physics suites (new or from other models) and DA systems be best integrated with FV³? How can they take advantage of FV³'s variable-resolution capabilities?
 - This is where the community can contribute the most!!
- What further improvements can be made for nesting and grid refinement?
- What features do physics/DA/ensembles developers need from FV³? Stochastic physics, IAU, adjoints, WAM, etc.
 - WAM Physics including horizontal diffusion and fast transport of major species; rapid response of thermosphere and ionosphere to space weather drivers; rapid updates to DA to resolve tides and meet space weather driver requirements
- ***Any changes to the dynamics would require close collaboration with FV³ core developers at GFDL and/or NASA-Goddard to ensure correctness of implementation.***